

UNITED STATES ENERGY RESEARCH AND DEVELOPMENT ADMINISTRATION CONTRACT W-7405-ENG. 36

Printed in the United States of America. Available from National Technical Information Service U.S. Department of Commerce 5285 Port Royal Road Springfield, VA 22161 Price: Printed Copy \$3.50 Microfiche \$2.25

,

٠

!

.

This report was prepared as an account of work sponsored by the United Niates Government. Neither the United Niates nor the United Niates Energy Research and Development Administration, nor any of their employees, nor any of their contractors, subcontractors, or their employees, makes any warranty, express or implied, or assumes any legal liability or responsibility for the accuracy, completeness, or usefulness of any information, apparatus, product, or process disclosed, or represents that its use would not infringe privately owned rights.

PRODUCTION OF A NARROW BAND OF 0.511-MeV RADIATION BY USE OF THE PHERMEX BREMSSTRAHLUNG SPECTRUM



Michael A. Stroscio

ABSTRACT

The pair production cross section is numerically integrated over a typical PHERMEX bremsstrahlung spectrum to obtain the probability of pair production in a target of nuclear charge Z, and density ρ . The pair production cross section used herein is only approximate in that it (1) neglects screening, (2) neglects the Coulomb field for the emerging pair (first Born approximation), and (3) neglects pair production by atomic electrons. In spite of these approximations, the present work still gives an order-ofmagnitude estimate of the amount of 0.511-MeV radiation produced by a typical pulse.

I. INTRODUCTION

Herein, the pair production cross section is numerically integrated over the PHERMEX output spectrum for a typical case given by Venable et al.¹ The pair production cross section for unpolarized photons of energy h ω is.^{2,3,4,5}

$$\begin{split} \sigma(\omega) &= \alpha Z^2 r_o^2 \left\{ 2\eta^2 [2C_2(\eta) - D_2(\eta)] \right. \\ &+ \frac{2}{27} \left[(109 + 64\eta^2) E_2(\eta) - (67 + 6\eta^2) (1 - \eta^2) F_2(\eta) \right] \right\}, \end{split}$$

where

$$C_2(\eta) = \int_1^{1/\eta} \frac{\cosh^{-1} x}{x} \cosh^{-1} \frac{1}{\eta x} dx \quad (\eta \le 1), \quad (2)$$

$$D_{2}(\eta) = \int_{1}^{1/\eta} \frac{\cosh^{-1} \frac{1}{\eta X}}{\sqrt{x^{2}-1}} dx \qquad (\eta \le 1), \quad (3)$$

$$E_{2}(\eta) = F(\sqrt{1-\eta^{2}}) - E(\sqrt{1-\eta^{2}}) \qquad (\eta \le 1), \quad (4)$$

 $F_2(\eta) = F(\sqrt{1-\eta^2})$ ($\eta \le 1$), (5)

and $\eta = 2mc^2/\hbar\omega$. In Eq. (1) α is the fine structure

constant, r_0 is the classical radius of the electron, and Z is the nuclear charge. F and E in Eqs. (4) and (5) denote the complete elliptic integrals of the first and second kind, respectively;

$$F(\sqrt{1-\eta^2}) = \int_0^{\pi/2} \sqrt{\frac{d\phi}{1-(1-\eta^2)\sin^2\phi}}$$
, and (6)

$$E(\sqrt{1-\eta^{2}}) = \int_{0}^{\pi/2} \sqrt{1-(1-\eta^{2})\sin^{2}\phi} \, d\phi \,.$$
 (7)

The basic integral over the PHERMEX bremsstrahlung output spectrum $P(\hbar\omega),$ is,

$$F = \int_{2mc^2}^{\hbar\omega max} P(\hbar\omega)\sigma(\hbar\omega)d(\hbar\omega), \qquad (8)$$

where ωmax is the maximum frequency contained in $P(\hbar\omega)$ and m is the electron rest mass.

II. INTEGRATION TECHNIQUES

The actual integrations involved in Eqs. (2) - (8) are completed by Gauss-Legendre integration algorithms. All integrals are written in the form,



$$I = \int_{-1}^{1} f(y) dy = \sum_{i=1}^{m} a^{i} (i) \{ f[y^{i}(i)] + f[-y^{i}(i)] \}, \quad (9)$$

where $a^{j}(i)$ and $y^{j}(i)$ are the Gauss-Legendre weights and coordinates, respectively.⁶

The weights $a^{j}(1)$ and the coordinates $y^{j}(1)$ are chosen so that Eq. (9) is <u>exact</u> when f(y) is a polynomial of degree 2m in y. All of the integrals involved in this calculation are accurately calculated by Gauss-Legendre sums with small values of m; i.e., all integrands are closely approximated by polynomials of low order.

III. POWER SPECTRUM

A typical power spectrum for the PHERMEX bremsstrahlung output¹ has been used in this work to estimate the value of Eq. (8). This normalized spectrum is defined by linear interpolation between the ordered pairs of energy and power spectrum intensity in Table I.

IV. PAIR PRODUCTION CODE

A FORTRAN code was written to evaluate Eq. (8) and the integrals in Eqs. (2) - (7). This program is capable of numerical integration of any function which is adequately approximated by a polynomial of degree 32 or less. In addition, any set of ordered pairs, as in Table I, is allowed in this code. The program, which is documented with comment cards, is listed in Table II and follows the notation of Eqs. (1) - (9). The final value calculated by the code must be multiplied by the factor $-\alpha Z^2 r_0^2$, which is contained in Eq. (1).

TABLE I

ENERGY SPECTRUM^a

E _i (MeV)	P _i
0.0	.220
6.0	.175
12.0	.145
24.0	.105
27.0	.075
28.5	.050
29.4	.000

^a(E₁, P₁) pairs represent the power spectrum used in evaluating Eq. (8). A typical power spectrum¹ was chosen for the present work.

V. PAIR PRODUCTION PROBABILITY

The probability that the normalized power spectrum will produce an electron-positron pair in the first millimeter of interaction with a target of atomic weight A, atomic number Z, and density ρ , is

$$P = F N_A \rho / A , \qquad (10)$$

where

$$F = -\alpha Z^2 r_0^2 \quad (Computer \ Output), \qquad (11)$$

and N_A is Avogadro's number. Equation (10), of course, does not include the effects of Compton scattering on the photons represented by $P(\omega)$. The magnitude of the Compton scattering cross section is, for many elements, comparable with the pair production cross section.⁵ However, it must be recalled that Compton scattering only redistributes the photon distribution and, thus, the only major influence on Eq. (1) is that some photons are scattered below the pair production threshold of $2mc^2$.

The Compton scattering contribution is relatively small for lead in a typical PHERMEX energy range and the pair production cross section dominates. Evaluating Eq. (10) for Pb, we find,

$$P = -\left(\frac{1}{137.}\right)(82)^2 (2.8 \times 10^{-13} \text{cm})^2 (-14.48)$$
$$\times \left(\frac{6.0225 \times 10^{23}}{207} \quad \frac{\text{atoms}}{\text{gram}}\right) \quad (11.35 \text{ g/cm}^3) \quad (12)$$
$$\stackrel{\sim}{\to} \quad 0.18/\text{mm}.$$

Thus the probability that the <u>unity-normalized</u> PHERMEX bremsstrahlung spectrum will produce a positron-electron pair is 0.18 for each mm of length of a Pb target. The normalization factor for the power spectrum (which should multiply Eq. (12) to give the number of positrons produced per mm) is given by the average number of photons per MeV in the PHERMEX output.

VI. PRODUCTION OF A NARROW BAND OF RADIATION FROM POSITRONS

Positrons produced by the above mechanism are stopped very quickly in Pb. This follows from (1) the Bethe stopping-power formula⁵ which indicates that an electron of 50 MeV has an average range of 12.5 mm in Pb, (2) the fact that maximum energy of the positrons considered herein is less than the maximum bremsstrahlung energy (29.4 MeV was taken as the maximum in Table I), and (3) the observation that high-energy positrons behave electromagnetically as high-energy electrons.

The cross section for the annihilation of an electron-positron pair into two photons is, 7,5

$$\sigma_{2\gamma} = \pi r_o^2 \frac{1}{\gamma + 1} \left[\frac{\gamma^2 + 4\gamma + 1}{\gamma^2 - 1} \ln \left(\gamma + \sqrt{\gamma^2 - 1} \right) - \frac{\gamma + 3}{\sqrt{\gamma^2 - 1}} \right]$$
(13)

where $\gamma = \frac{1}{\sqrt{1-\beta^2}}$, $\beta = v/c$, and v is the velocity of the positron with respect to the electron at rest. Using the relations

$$\sqrt{\gamma^2 - 1} = \gamma \beta \tag{14a}$$

and

$$\ln \left(\gamma + \sqrt{\gamma^2 - 1}\right) = \frac{1}{2} \ln \left(\frac{1 + \beta}{1 - \beta}\right) , \qquad (14b)$$

and expanding the logarithm in Eq. (14b) we have,

$$\sigma_{2\gamma} = \pi r_{o}^{2} \frac{1}{\gamma+1} \left[\frac{\gamma+1}{\gamma^{2}\beta^{2}} \beta + \frac{\gamma_{2}^{4}\beta\gamma+1}{\gamma^{2}\beta^{2}} \left(\frac{1}{3} \beta^{3} + \frac{1}{5} \beta^{5} + \cdots \right) \right]$$

$$= \frac{\pi r_{o}^{2}}{\gamma^{2}\beta} + \pi r_{o}^{2} \frac{\gamma^{2}+4\gamma+1}{\gamma^{2}(\gamma+1)} \left(\frac{\tanh^{-1}\beta-\beta}{\beta^{2}} \right) \qquad (15)$$

$$= \frac{\pi r_{o}^{2}}{\gamma^{2}\beta} + \pi r_{o}^{2} \frac{\gamma^{2}+4\gamma+1}{\gamma^{2}(\gamma+1)} \left(\frac{\beta}{3} + \frac{\beta^{3}}{5} \frac{\beta^{5}}{7} + \cdots \right) \qquad .$$

Upon expanding γ in Eq. (15) in terms of β we find,

$$\sigma_{2\gamma} = \frac{\pi r_0^2}{\beta} - \pi r_0^2 \beta + \pi r_0^2 \beta + \pi r_0^2 \beta + \pi r_0^2 \beta + \pi r_0^2 \left(\frac{3}{5} - \frac{3}{4}\right) \beta^3 + 0 \ (\beta^5)$$

$$= \frac{\pi r_0^2}{\beta} - \pi r_0^2 \frac{3}{20} \beta^3 + 0 \ (\beta^5) \ ; \qquad (16)$$

i.e., the linear terms in β cancel. Equation (16) indicates that the major contribution to the twophoton annihilation of an electron-positron pair occurs for small values of β . Thus the kinetic energies of the annihilating particles are small compared to their rest mass. This implies that (1) annihilation will result in two photons of approximately 0.511-MeV energy in opposite directions and (2) the distribution of this narrow band of 0.511-MeV radiation will be isotropic allowing for any observation angle that suits the experimenter. The onephoton annihilation is, of course, forbidden by conservation of angular momentum. The three-photon cross section is smaller than the two-photon cross section of Eq. (13) by a factor of (1/137). Ref. 8 contains a review of the theoretical annihilation characteristics of electron-positron pairs for all order processes which have been observed or are likely to be observed for some time.

VII. DISCUSSION

As shown above, most of the photons in $P(\omega)$ will produce a positron if the target is several millimeters thick. This means that a large fraction of the energy in $P(\omega)$ will appear as 0.511-MeV radiation in a narrow band.

This narrow band of high-intensity radiation has not been exploited for any useful purpose at the PHERMEX facility. Among the various uses of this radiation are (1) the measurement of opacities at 0.511 MeV and (2) the measurement of solid state properties from a study of the exact annihilation spectrum. This last use has received considerable attention⁹ and is commonly used to determine the solid state properties of the material in which the positrons are produced. The unique feature here is that these measurements would be made in the presence of an intense bremsstrahlung spectrum.

It is clear that not all of the 0.511-MeV γ rays produced in the Pb target will escape without interacting with the Pb target itself. At 0.511 MeV, γ rays interact with Pb by both Compton scattering and by the photoelectric effect.⁵ The attenuation of 0.511-MeV radiation in Pb is described by an exponentially decreasing intensity, I:

$$I = I_o e^{-TX}$$
,

where I_o is the intensity at the point of production, x is the distance through which the γ rays travel in Pb and $\tau = 0.17$ per mm.⁵ Thus the intensity of 0.511-MeV γ rays is diminished by the factor e^{-1} in about 6 mm of Pb. In comparison, an electron with 5 MeV (50 MeV) of kinetic energy has a range⁵ of 3.3 mm (12.5 mm).

These data indicate that the optimum 0.511-MeV pulse will be obtained when some dimension of the Pb target is restricted to about 5 mm in extent: this thickness of Pb will stop most of the positrons produced in the PHERMEX energy range and will allow about 50% of the 0.511-MeV radiation produced by positron annihilation to escape unattenuated. A particularly attractive target design consists of an array of cylinders of Pb, each about 5 mm in diameter and 50 to 100 mm long, with their axes parallel to the PHERMEX beam. This target would provide a long interaction distance for the PHERMEX photons and would allow most of the 0.511-MeV radiation produced at 90° to the incident beam to escape the Pb.

ACKNOWLEDGMENT

The author is especially grateful to John W. Taylor, LASL Group M-2, for several stimulating discussions regarding the production of positrons at the PHERMEX facility and for providing the power spectral intensity of the PHERMEX x-ray beam which was used in this work.

REFERENCES

- D. Venable, D. O. Dickman, J. N. Hardwick,
 E. D. Bush, Jr., R. W. Taylor, T. J. Boyd,
 J. R. Rube, E. J. Schneider, B. T. Rogers, and
 H. G. Worstell, "PHERMEX: A Pulsed High Energy Radiographic Machine Emitting X-rays," LA-3241, (1967).
- R. Jost, J. M. Luttinger, and M. Slotnick, Phys. Rev. <u>80</u>, 1950, 189.
- F. Rohrlich and R. L. Gluckstern, Phys. Rev. <u>86</u>, 1957, 189.
- J. M. Jauch and F. Rohrlich, <u>The Theory of Photons and Electrons</u> (Addison-Wesley Publ. Co., <u>Reading, MA</u>, 1955) p. 376.
- W. Heitler, <u>The Quantum Theory of Radiation</u>, 3rd Ed. (Oxford at the Clarendon Press, Glasgow, 1954).
- F. Schied, <u>Numerical Analysis</u> (McGraw-Hill Book Co., New York, 1968) p. 125 et seq.
- P. A. M. Dirac, Proc. Camb. Phil. Soc. <u>26</u>, (1930)
 p. 361.
- M. A. Stroscio, Physics Reports, <u>22C</u>, (1975), p. 217.
- A. T. Stewart, L. O. Roelig, editors, <u>Positron</u> <u>Annihilation</u> (Academic Press, 1967) Proceedings of the June 1965 Positron Annihilation Conference at Wayne State University, Detroit, MI.

TABLE II

PAIR PRODUCTION CODE^a

(LASL Identification: LP-0640)

```
PROGRAM PUBLIS(INP+FSFT5=INP+OUT+FSFT6=OUT)
       DIMENSION Y(11,16) +A(11,16) +IM(11)
       COMMON Y.A
       NORD=11
          HERE WE READ THE PARAMETERS FOR THE GAUSS-LEGENDRE
C
          INTEGRATION POUTINE
C
       DO I MM=1+NORD
       RFAD(5,100) N
 100 FORMAT(110)
       IND=N/2
       TM(MM) = TND
       RFAD(5,101) (Y(MM ,1),A(MM ,1),I=1,IND)
  101 FORMAT(RE10.8)
    1 CONTINUE
       M=4
       SIJM±0_0
       IND=IM(M)
          HERE WE INTEGRATE THE PAIP PROD. CROSS SECTION OVER THE POWER SPECTRUM OF THE INITIAL PHOTON DISTRIBUTION
C
C
       DO 3 1=1.1ND
       SUM=SIM+A(M,I)*(F(Y(M,I)) + F(-Y(M,I)))
    3 CONTINUE
       NPOINT=IM(M)*2
       WRITE(6,102) NPOINT,SUM
  102 FORMAT(1X, 14, 3X, F15, 8)
       STOD
       END
       FUNCTION F(Y)
          OFMAX IS THE MAXIMUM EREQUENCY OF THE PHOTON DISTRIBUTION
C
          EN IS THE ELECTRON REST MASS IN MEV
C
          A*Y+B IS THE FREQUENCY OF THE PHOTON DISTRIBUTION
P REPRESENTS THE POWER DISTRIBUTION OF INITIAL PHPTONS
c
c
C
          SIGMA IS THE TOTAL CROSS SECTION FOR PAIR PRODUCTION
       OFMAX=29.4
       FM= 511002
       \Delta = (OFM\Delta X - 2 \bullet FM) / 2 \bullet
       P=( 1FMAX+7 . #FM1 / 7 .
       F=A+P(A+Y+R)+SIGMA(2.+FM/(A+Y+R))
       WOTTFIN,441 F
      FORMAT(1X+#F=#+F15.8)
  44
       RETURN
       FND
       FUNCTION SIGNA(FTA)
          FTA IS 2M/OMEGA AND IS ALWAYS SMALLER THAN 1 FOR PAIR PROD.
C
       STGMA=-2.*(FTA++2)*(2.*C2(FTA)-D2(FTA))
      1 -(2./27.)*((109.+64.*(FTA)**2)*(FA(ETA)-FA(ETA))
          -(67++6+*(FTA)**2)*(1+-FTA**2)*FA(FTA))
      2
       OMEGA1=1.02/ETA
       WRITE(6+492) SIGMA+FTA+OMEGA1
 492 FORMAT(1X,**SIGMA=**F15.8,3X,**FTA=**F15.8,3X,**OMFGA=**E15.8)
       RETURN
       END
```

^aFORTRAN code for integrating the power spectrum over the pair production cross section.

```
DIMENSION F(7), ST(7)
      THE INITIAL PHOTON POWER SPECTRUM IS READ IN AS SEVEN
      PATRS OF NUMBERS AND THE BELOW ROUTINE DOES A LINEAR FIT
      TO THESE SEVEN ORDERED PAIRS
                                            THE ENERGIES E(1) THRU
      E(7) ARE ORDFRED WITH THE SEVEN INTENSITIES SI(1) THRU SI(7)
   F(1)=0.0
   F(2)=6.0
   F(3)=12.0
   F(4)=24.0
   F(5)=27.0
   F(A)=28.5
   F(7)=20.4
   St(1)=.220
   ST(2)=.175
   51(3)=.145
   St(4)=+105
   ST(5)=.075
   ST(6)=.050
   ST(7)=.000
   IF (CFRFQ.GT.F(1)) K=1
   IF (CFRFQ.GT.F(2)) K=2
   IF (CEREQ.GT.F(3)) K=3
   IF (CFRFQ.GT.F(4)) K=4
   IF (CFRFQ.GT.F(F)) K=5
   IF (CEPEQ.GT.F(A)) K=6
   P = SI(K) + ((SI(K+1) - SI(K))/(F(K+1) - F(K))) * (CFRFQ - F(K))
   RETURN
   END
   FUNCTION COLFTAN
   DIMENSION Y(11,16), A(11,16)
   COMMON Y,A
   COMMON /FRT1/ AAA
   AAA = FTA
   501=0.0
   MM=16
   NN=11
   DO 2 J=1,MM
   SU1=Sij]+A(NN+J)*(F1(Y(NN+J)) + F1(-Y(NN+J)))
 2 CONTINUE
   C2=5U1
   WPTTE(6+50) C2
50 FORMAT(1X++C2=++F15+8)
   RETURN
   END
   FUNCTION FI (11)
   COMMON /FRT1/ AAA
   FTA=AAA
   F1=ACOSH(.5*(1./FTA-1.)*U+.5*(1.+1./ETA))
   1 *ACOSH(1./(.5*(1.-ETA)*U+.5*(ETA+1.)))
     *(1./(U+(1.+1./FTA)/(1./FTA-1.)))
   2
   RETURN
   FND
   FUNCTION D2(FTA)
   DIMENSION Y(11,16), A(11,16)
   COMMON Y.A
   COMMON /FRT2/ RPR
   BBREFTA
    SU1=0.0
   MM=16
    NN=11
   DO 2 J=1,MM
    SU1 = SU1 + A(NN + J) + (F2(Y(NN + J)) + F2(-Y(NN + J)))
 2 CONTINUE
   D2=5U1
   WRITE(6,59) D2
59 FORMAT(1X+*D2=*+E15+8)
    RETURN
    FND
```

· -- |

1

١.

c c c c

FUNCTION P(CEPEO)

```
FUNCTION F2(U)
    COMMON /FRT2/ BRR
    FTA=RRR
    F2=((1.-FTA)/(2.*ETA))*
   1 SQRT(1./((.5*(1./FTA-1.)*U+.5*(].+1./FTA))**2-1.))
       *ACOSH(1./(.5*(1.-FTA)*U+.5*(FTA+1.)))
   2
    RETURN
    FND
    FUNCTION FA(FTA)
    DIMENSION Y(11+16)+A(11+16)
    COMMON Y.A
    COMMON /FRT3/ CCC
    CCC=FTA
    SU1=0.0
    MM=16
    NN=11
    DO 2 J=1+MM
    SU1=SU1+A(NN+J)*(F3(Y(NN+J)) + F3(-Y(NN+J)))
  2 CONTINUE
    FA=SU1
    WRITE(6+62) FA
    FORMAT(1X,*F=*,F15.8)
62
    RFTURN
    END
    FUNCTION FR(I)
    COMMON /FRT3/ CCC
    FTA=CCC
    PTOF=3.1415926538/4.
    ACRGTP= (1.-FTA##2)
F3=PIOF# SQRT(1./(1.-(ACBGTR
                                    )*SIN(PIOF*(1.+U))
   1 *SIN(PIOF*(1.+U))))
    RETURN
    FND
    FUNCTION FA(FTA)
    DIMENSION Y(11,16) + A(11,16)
    COMMON Y .A
    COMMON /FRT4/ DDD
    DDD=FTA
    SU1=0.0
    MM=16
    NN = 1T
    00 2 J=1,MM
    SU1=SU1+A(NN+J)*(F4(Y(NN+J)) + F4(-Y(NN+J)))
  2 CONTINUE
    FA=SU1
    WRITE(6,78) FA
78 FORMAT(1X++F=++F15+R)
    RETURN
    END
    FUNCTION F4(11)
    COMMON /FRT4/ DDD
    FTA=DDD
    PIOF=== 1415926538/4.
    ACRGTR= (1.-FTA##2)
    F4=PIOF#SORT(].-(ACBGTR
                              ) #SIN(PIOF#(1.+U))
     *SIN(PIOF*(1.+U)))
   1
    RETURN
    FND
    FUNCTION ACOSH(XX)
    ACOSH=ALOG(XX+SORT(XX**2-1.))
    RETURN
    FND
```

```
$FM.
```

2			
.57735027 1.nnnnn	1000		
4			
•86113631 • 347854	85 .33998104 .65214515		
6			
•93246951 •171324	49 •66120939 •36076157	•23861919 •46791393	
8			
•96028986 •101228	54 •79666648 •22238103	•52553241 •31370665	•18343464 •36268378
10			
•97390653 •066671	34 •86506337 •14945135	•67940957 •21908636	•43339539 •26926672
•14887434 •295524	22		
12			
•98156063 •047175	•34 •0411726 •10693933	•76990267 •16007833	•58731795 •20316743
•36783150 •233492	254 •12523341 •24914705		
14			
•98628381 •n35119	46 • • 2843488 • 08015809	•B2720131 •12151857	•68729290 •15720317
•51524864 •185538	40 •31911237 •20519846	•10805495 •21526385	
16			75510111 10110007
•98940093 •n27152	46 •94457502 •06225352	•86563120 •09515851	•/5540441 •12462897
•01/0/024 •149595	•45801678 •16915652	•20100355 •10200341	•09501251 •16945061
		27270(00 1/200(1)	5100(700 131/00//
•V/022022 •192/23	537 + 2277 + 235 + 14917299	• 37370609 • 14209611	01222663 06267205
•03003300 •[10194		• 0 3 9 1 1 6 9 7 • 0 0 3 2 7 6 7 4	•91223443 •08287203
•70371193 •(140001	45 • 49912000 • 01701401		
74 06405680 197039	10111007 10593746	21504260 12167047	. 42270251 . 11550547
6/6/01/07 107///	37 44809346 00741866	74012410 08410014	- 43574551 - 1550567 - 82000100 - 07234448
89661552 050209	121 004007707 017701000 159 02827455 04437744	07/72954 02952120	00519732 01234122
22	JJC 8-30214JJ 804421144	• 714120 <u>10</u> • 020331.19	• 4 9 5 1 8 7 2 2 • (1 2) 4 1 2 5
.04830767 .096540	09 .14447196 .09563872	-23928736 -09384440	-33186860 -09117388
42135128 087652	209 50689991 08331192	-58771576 -07819390	•66304427 •07234579
-73218212 -065822	22 .79448380 .05868409	84936761 05099806	•89632116 •04283590
•9349060B •034273	386 .96476226 .02539207	•98561151 •01627439	•99726386 •00701861
SFJ.			