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A Quantum- and Correlation-Corrected
Thomas-Fermi-Dirac Equation
with a FORTRAN Code



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### LOS ALAMOS SCIENTIFIC LABORATORY LOS ALAMOS of the New MEXICO University of California

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# A Quantum- and Correlation-Corrected Thomas-Fermi-Dirac Equation with a FORTRAN Code

by

John F. Barnes





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### ABSTRACT

In a previously published report (IA-2750, A Proposed Modification of the Available Momentum Region in Thomas-Fermi Theory) a "quantum correction" to the statistical model of the atom was derived in detail. In the present work we use these results and a very simple approximation to the correlation energy in deriving a quantum— and correlation—corrected Thomas-Fermi-Dirac (TFD) equation. One expects the radial density distribution and potential calculated from this equation to be improved over those on the TFD model, both near the nucleus and near the outer boundary of the atom or ion. Minimum—energy (that is, zero boundary pressure) solutions for rare—gas atoms possess values of cell radius in good agreement with those calculated from experimental values of the lattice parameter.

A FORTRAN code is included in an Appendix to the report.

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### I. INTRODUCTION

Many properties of free atoms and of solids are predicted with considerable accuracy by the statistical atom model of Thomas, Fermi, and Dirac. 1-4 The accuracy of the method has probably helped inspire the deeper investigation into its foundations; for if one considers only the premises upon which the original model is based, he would hardly expect more than very rough agreement with experiment or with the predictions of a more refined theory. Yet, to give some examples, the radial density distribution function for a given element calculated from either the Thomas-Fermi (TF) or the Thomas-Fermi-Dirac (TFD) equation appears to be, for any but the very lightest elements, in good agreement with that obtained from self-consistent field calculations. The shell structure is not reproduced, but there appears to be a reasonable averaging of the relative maxima and minima. One-electron energy levels computed in the TF or TFD fields are remarkably close to experimental term values, and the atomic number at which electrons of a given angular momentum make their "first appearance" are predicted correctly to the nearest integer.

The correctness of these predictions would have to be regarded as fortuitous, when they result from a theory which presumes that (1) the potential field in an atom varies sufficiently slowly that the fractional change in an electron's de Broglie wave length  $\lambda$  is small over a distance comparable to  $\lambda$ ; and (2) there are a great many electrons in a region having a volume of the order of  $\lambda^3$ . However, Fényes showed that the TF energy and density expressions follow from quasi-classical arguments as well as from a strictly classical viewpoint. Later, March and Plaskett succeeded in deriving the TF energy equation from an integral of the

WKB eigenvalues over a particular region of the quantum-number plane. Thus, the validity of the statistical model seems to depend upon that of the WKB approximation, and Langer showed that the WKB phase integral, at least in one dimension, is applicable under quite broad conditions on the potential even for small quantum numbers.

Since it is known that the statistical theory rests on more than the classical arguments, it appears well worthwhile to attempt those improvements that can be made with little complication of the equations to be solved. The derivation of March and Plaskett provides the basis for a "quantum correction" of the statistical model reported earlier 8,9 and outlined below.

We have stressed the general success of the TF and TFD equations. There are, nevertheless, some areas in which the calculations do not agree well with experiment. A very apparent discrepancy is in the total binding energy of the electron cloud. The electron density predicted at the nucleus is infinite on either the TF or the TFD model; consequently, the calculated binding energy is considerably too large. The previously derived quantum correction modifies the density in the region near the nucleus and produces binding energies in much better agreement with experimental values and with those obtained from self-consistent field calculations.

A further area for attempting improvement is suggested by the rather large errors in the calculated pressure-compression curves. These relationships are influenced mainly by the outermost parts of the electron distribution. In this region of low density the correlation energy, neglected in the original models, becomes important.

In the present work we incorporate the quantum correction and a very simple form for the correlation energy in the derivation of an easily applied quantum- and correlation-corrected TFD equation. One expects the density distribution and potential calculated from this equation to be improved over those on the TFD model, both near the nucleus and near the outer boundary of the atom or ion.

The numerical procedures used in obtaining solutions of the equation are discussed, and a few calculational results are summarized. It is found that the inclusion of the correlation energy should not greatly change the pressure-compression curves for most elements. However, minimum-energy (that is, zero boundary pressure) solutions for rare-gas atoms possess values of cell radius in good agreement with those calculated from experimental values of the lattice parameter. These results suggest the interpretation of minimum-energy solutions as representing isolated atoms, rather than atoms in crystals, since the rare gases are known to be bound in crystals by the very weak van der Waals forces.

Numerical work was performed on an IHM 7030 computer. A FORTRAN code, version "F4", is listed in the Appendix, but we should caution that certain changes, mainly in the Input-Output statements, might have to be made before using the code with other computer systems. Also, it is necessary to carry more than eight-figure precision throughout the calculations in order to obtain accurately the solutions possessing mimimum energy. We have attempted to carry our calculations to about thirteen figures, and while the 7030 word size is equivalent to about 16 decimal digits, it would be necessary to perform the calculations in double-precision arithmetic on a smaller machine. Eight-figure precision is certainly sufficient, however, to obtain general solutions corresponding to arbitrary degrees of compression.

### II. A QUANTUM- AND CORRELATION-CORRECTED TFD EQUATION

### A. The TF and TFD Equations

A variational technique can be used to derive the TF equation, and an extension of this method provides an often-used and simple means of adding corrections to the statistical model. Thus, we can write the Fermi kinetic energy density of a gas of free electrons at a temperature of zero degrees absolute in the form

$$v_{\rm p} = c_{\rm p} p^{5/3},$$

where

$$c_{p} = (3/10)(3\pi^{2})^{2/3}.*$$

The electrostatic potential energy density is the sum of the electronnuclear and the electron-electron terms. We can write this as

$$U_{p} = U_{p}^{n} + U_{p}^{e} = -(v^{n}+v^{e}/2)_{\rho},$$

where  $v^n$  is the potential due to the nucleus of charge Z;  $v^e$  is the potential due to the electrons; and the factor of 1/2 is included in the electron-electron term to avoid counting each pair of electrons twice. With x denoting distance from the nucleus, the total energy of the spherical distribution is given by

$$E = \int [c_{f} \rho^{5/3} - (v^{n} + v^{e}/2) \rho]^{4\pi x^{2}} dx.$$
 (1)

The expression for density on the TF model,

$$\rho = \sigma_0(E'-V)^{3/2}, \qquad (2)$$

with

$$\sigma_0 = (3/5c_f)^{3/2}$$

is obtained by minimizing Eq. (1) subject to the auxiliary condition that the total number of particles, N, remains constant. The potential

Throughout this work we shall use atomic units (a.u.), in which e = h = m = 1. The unit of length is  $a_0$ , the first Bohr radius for hydrogen; and the unit of energy is  $e^2/a_0$ .

energy, V, is a function of position in the electron distribution; E' is the Fermi energy, or chemical potential, and is constant throughout a given distribution. The TF equation follows from Eq. (2) and Poisson's equation.

The tendency for electrons of like spin to stay apart because of the exclusion principle is accounted for by the inclusion in Eq. (1) of the exchange energy, the volume density of which is given by

$$U_{\rm ex} = -c_{\rm ex}\rho^{4/3},$$

where

$$c_{ex} = (3/4)(3/\pi)^{1/3}$$
.

Minimization of the total energy now leads to the equation

$$(5/3)c_{\rm f}\rho^{2/3} - (4/3)c_{\rm ex}\rho^{1/3} - (E'-V) = 0,$$

which is quadratic in  $\rho^{1/3}$ . From this equation we get

$$\rho = \sigma_0 [\tau_0 + (E' - V + \tau_0^2)^{1/2}]^3, \tag{3}$$

where

$$\tau_0 = (4c_{\rm ex}^2/15c_{\rm f})^{1/2}$$
.

Now Poisson's equation with the density given by Eq. (3) leads to the TFD equation.

In the following two sections we propose additional energy terms to be included in Eq. (1); the incorporation of these terms leads to a simple quantum— and correlation—corrected TFD equation.

### B. The Quantum Correction

The quantum-correction energy density follows from a slight change in the derivation due to March and Plaskett; the development will be only briefly outlined here.

March and Plaskett have demonstrated that the TF approximation to the sum of one-electron eigenvalues in a spherically symmetric potential is given by the integral

$$I = 2 \int \int (2\ell+1)E(n_r,\ell)dn_rd\ell, \qquad (4)$$

where the number of states over which the sum is carried is written as

$$N = 2 \int \int (2l+1) dn_{\mathbf{r}} dl.$$
 (5)

Here  $E(n_r,\ell)$  is the expression for the WKB eigenvalues considered as functions of continuous variables;  $n_r$  is the radial quantum number;  $\ell$  is the orbital quantum number; and the region of integration is bounded by  $n_r = -1/2$ ,  $\ell = -1/2$ , and  $E(n_r,\ell) = E'$ . We have included a factor of two in these equations to account for the spin degeneracy of the electronic states. The Fermi energy E' is chosen so that Eq. (5) gives the total number of states being considered, the N electrons occupying the N lowest states. With considerable manipulation, Eq. (4) becomes the TF energy equation

$$I = \int \left(\frac{3}{5} \frac{p^2}{2} + V\right) \frac{p^3}{3\pi^2} 4\pi x^2 dx, \qquad (6)$$

and Eq. (5) reveals the TF density through the expression

$$N = \int \frac{P^3}{3\pi^2} 4\pi x^2 dx, \qquad (7)$$

both integrals being taken between the roots of E' = V(x). We have written these results in atomic units, so that P, the Fermi momentum,

is defined by

$$P = 2^{1/2} (E'-V)^{1/2}.$$
 (8)

It is pertinent to examine the error in the TF sum of eigenvalues, as given by Eq. (6), for the case of the pure Coulomb field. The WKB eigenvalues in a Coulomb field are given by

$$E_{n_r,\ell} = -Z^2/2(n_r+\ell+1)^2,$$

and let us consider the levels filled from n = 1 to n = v, where n is the total quantum number defined by

$$n = n_r + \ell + 1.$$

Then, for any value of  $\nu$  we can evaluate the error in the TF approximation to the sum of eigenvalues, comparing always with the correct value,  $-z^2\nu$ . Scott's correction to the total binding energy is obtained by letting  $\nu$  become very large.

Although the sum of one-electron eigenvalues is not the total energy of the statistical atom because of the electron-electron interaction being counted twice, we might expect to improve the calculated binding energy greatly by correcting this sum in some manner, since the chief cause of the discrepancy is certainly the large error in the electron-nuclear potential energy. This correction can be performed by imposing a new lower limit on  $\ell$  in the integrations above. When we introduce a new lower limit  $\ell_{\min}$  and a related quantity which we call the "modification factor,"

$$a = \ell_{\min} + 1/2,$$

we obtain, after more manipulation, slightly different expressions corre-

sponding to Eqs. (6) and (7). From these revised expressions we can identify a quantum-corrected TF density expression,

$$\rho = \sigma_0 (E' - V - a^2 / 2x^2)^{3/2}, \qquad (9)$$

and a corrected kinetic energy density,

$$U_{k} = c_{f} \rho^{5/3} + (a^{2}/2x^{2})\rho. \tag{10}$$

The revised lower limit on the volume integrals, say  $x_1$ , is the lower root\* of

$$E' - V - a^2/2x^2 = 0;$$
 (11)

for  $x < x_1$ ,  $\rho$  must vanish, and we have thus termed  $x_1$  the "inner density cutoff distance." We can call the second term on the right-hand side of Eq. (10) the "quantum-correction energy density" and write it in the more consistent form

$$U_{q} = (c_{q}/x^{2})\rho, \qquad (12)$$

by defining

$$c_0 = a^2/2$$
.

The modification factor, a, is determined by the initial slope of the potential function, as described in Part III of this report.

For interpreting these results it is helpful to consider just what we have done in changing the lower limit of the orbital quantum number.

<sup>\*</sup>In application to the atomic problem, there is only one root of Eq. (11) between zero and the outer boundary of the atom or ion. This root is identified as  $x_1$ , and  $x_2$  is then determined by the usual TF boundary condition.

Since the lower limit  $\ell=-1/2$  must correspond to an orbital angular momentum of zero, we have, clearly, eliminated states with angular momentum of magnitude between zero and a cutoff value  $L_c=a\hbar$ . Corresponding to  $L_c$  at every radial distance is now a linear cutoff momentum

$$p_c = a\hbar/x$$
,

and we can rewrite Eq. (9) in terms of the Fermi momentum and cutoff momentum:

$$\rho = (\sigma_0/2^{3/2})(P^2 - p_c^2)^{3/2}.$$

At radial distances less than  $x_1$ , momenta are prohibited over the entire range from zero to P, so the electron density vanishes.

This interpretation must be modified somewhat when exchange and correlation effects are included; for then the Fermi momentum is no longer simply given by Eq. (8), except very near the nucleus. We can define  $\mathbf{x}_1$  as in the absence of interactions, i.e., as the lower of the roots of Eq. (11), but it is not correct to demand that the density vanish at the upper root. Instead, we require only that the density be real.

### C. The Correlation Correction

The original TF equation describes a system of independent\* particles, while the introduction of exchange energy, which leads to the TFD equation, represents a correction for the correlated motion of electrons of like spin. The remainder of the energy of the electron gas is termed the correlation energy; by its inclusion we are recognizing that electrons, regardless of spin orientation, tend to avoid one another.

The particles are "independent" in the sense that there is no correlation among their positions. They do interact with each other, however, in establishing the potential field in which each particle moves.

In extensions of the statistical model there have been suggested at least two different expressions  $^{11,12}$  for the correlation energy that approach, in the appropriate limits, Wigner's low-density formula and the expression due to Gell-Mann and Brueckner at high densities. In addition to these,  ${\tt Gombás}^{13}$  and  ${\tt Tomishima}^{14}$  have utilized expansions of the correlation energy per particle in powers of  $\rho^{1/3}$  about the particle density encountered at the outer boundary of the atom or ion. In this expansion, the term of first-order can be considered as a correction to the exchange energy, and it follows that the TFD solutions for a given Z then correspond to correlation-corrected solutions for a modified value of Z. Aside from the rather poor approximation of the correlation energy, a drawback to this procedure is that the TFD solutions must be at hand. If solutions representing specified degrees of compression are desired, the method would appear to be impractical.

It is, however, interesting and fortunate that over the density range of interest it is apparently possible to approximate the correlation energy per particle quite closely by an expression of the form

$$u_c = -c_c \rho^{1/6}$$
. (13)

This is shown in Fig. 1, where we have set  $c_c = 0.0842$ , and compared this approximation with the values due to Carr and Maradudin. The latter are obtained as a higher-order correction to Gell-Mann and Brueck-ner's formula at the high densities (say  $\rho \ge 0.25$  a.u.), and are again reasonable interpolated values at the lower electron densities.

There is no need to be concerned with the correlation energy outside the limited range of density shown in Fig. 1. The lowest density that can be obtained in solutions of the "corrected" TFD equation to be derived is about 0.002 a.u., and at densities above 1.0 a.u. the correlation energy becomes small compared with the exchange energy.

Near the lower limit of density, the magnitude of the correlation energy computed from Eq. (13) is about one-third as large as the exchange energy,

Figure 1. Correlation energy per electron.

but at  $\rho=2.0$  a.u. it is only 10% as large. For, say,  $\rho=10^5$  the ratio is 1% and the exchange energy itself is only 0.01% of the Fermi kinetic energy.

We shall, then, approximate the correlation energy density with

$$U_c = -c_c \rho^{7/6}$$

and, in atomic units,

$$c_2 = 0.0842.$$

### D. Derivation of the Equation

From the results of the preceding paragraphs, we can now express the total energy per unit volume of the charge distribution in the form

$$U = c_f^{5/3} - c_{ex}^{4/3} - c_c^{7/6} - (v^n + v^e/2)\rho + (c_q/x^2)\rho,$$

where all quantities appearing in the equation have been previously defined. By minimizing the integral of U over the volume occupied by the charge, while requiring that the total number of electrons be fixed, we obtain the following equation:

$$\rho^{2/3} - \tau_1 \rho^{1/3} - \nu_0 \rho^{1/6} - R/4 = 0, \qquad (14)$$

where

$$\tau_1 = (4/5)(c_{ex}/c_f),$$

$$v_0 = (7/6)c_e^{\sigma_0^{2/3}},$$

$$R = 4\sigma_0^{2/3}(E'-V-c_q/x^2).$$

The electron density is found as a function of R by solving Eq. (14), a quartic in  $\rho^{1/6}$ . To accomplish this we write a "resolvent cubic equation" in terms of another variable, say y:

$$y^3 + \tau_1 y^2 + Ry + (\tau_1 R - v_0^2) = 0.$$
 (15)

Let us use the same symbol, y, to denote any real root of this cubic equation. We can then express the four roots of the quartic, and hence four expressions for the electron density, in terms of y. One of these expressions possesses the proper behavior in reducing to previously obtained results in the neglect of correlation and exchange effects, namely,

$$\rho = (1/8)[\tau_1 + \psi + (y^2 + R)^{1/2}]^{3}, \tag{16}$$

where

$$\psi = (\tau_1 + y)^{1/2} [\tau_1 - y + 2(y^2 + R)^{1/2}]^{1/2}. \tag{17}$$

We note that  $\psi$  vanishes when correlation is neglected, since  $y = -\tau_1$  is then a root of Eq. (15).

In the familiar manner we now define a modified TFD potential function  $\phi$  by the relation

$$Z\phi = (E'-V+\tau_0^2)x, \qquad (18)$$

and from Poisson's equation and Eq. (16) we obtain

$$\phi'' = (\pi x/2Z) [\tau_1 + \psi + (y^2 + R)^{1/2}]^3, \quad x \ge x_1,$$

$$= 0, \quad x < x_1. \quad (19)$$

In terms of  $\phi$ ,

$$R = 4\sigma_0^{2/3} (z\phi/x - a^2/2x^2 - \tau_0^2).$$
 (20)

Eqs. (20), (15), (17), and (19) constitute the differential relationship to be satisfied at each step in the integration. We could, of course, write immediately the solutions of Eq. (15) in analytic form, but it proves convenient in the numerical treatment to obtain a root by the Newton-Raphson method, since a good first guess in the iteration is available from the previous integration step.

The boundary conditions on Eq. (19) are: (1) As the nucleus is approached the potential must become that of the nucleus alone, or

$$\phi(0) = 1$$

and (2) at the outer boundary,  $x_2$ , of the distribution of N electrons,

$$N = \int_{x_1}^{x_2} \rho 4\pi x^2 dx$$

$$= Z \int_{\mathbf{x}_{1}}^{\mathbf{x}_{2}} \phi'' \mathbf{x} d\mathbf{x}.$$

Integration by parts yields

$$(\phi'x-\phi)_{x_{1}}^{x_{2}} = N/Z,$$

and since

$$\phi(x_1) = 1 + x_1 \phi'(x_1),$$

we have the usual condition:

$$\phi(x_2) = x_2 \phi'(x_2) + (Z-N)/Z.$$
 (21)

In addition to potential and density distributions, total binding energies of atoms are of special interest to us here. For the proper evaluation of energies, the arbitrary constant that is present originally in both the electrostatic potential energy and the Fermi energy must be specified. The state of infinite separation of the constituent particles is normally taken to have zero energy; we therefore follow the usual convention and fix the potential at the edge of the neutral atom at zero for all values of  $x_2$ . For an ion the potential energy of an electron at the boundary is taken as

$$V = -(Z-N)/x_2.$$

The defining relation, Eq. (18), now gives at the boundary

$$Z\phi(x_2) = [E' + (Z-N)/x_2 + \tau_0^2]x_2$$

or, solving for the Fermi energy,

$$E' = Z\phi(x_2)/x_2 - (Z-N)/x_2 - \tau_0^2$$

The total electron-nuclear potential energy is given by

$$E_p^n = -\int_{x_1}^{x_2} (z/x) \rho^{4\pi x^2} dx,$$

while for the electron-electron potential energy we have

$$E_p^e = (1/2) \int_{x_1}^{x_2} v^e \rho^{4\pi x^2} dx.$$

From Eq. (18) and the relation  $V = -(v^n + v^e)$ , this becomes

$$E_p^e = (1/2) \left[ -E_p^n + \tau_0^2 N + E'N - \int_{x_1}^{x_2} (Z\phi/x) \rho^{4\pi x^2} dx \right].$$

Other energy integrals are, with an obvious notation,

$$E_{f} = c_{f} \int \rho^{5/3} 4\pi x^{2} dx,$$

$$E_{q} = c_{q} \int (\rho/x^{2}) 4\pi x^{2} dx,$$

$$E_{ex} = -c_{ex} \int \rho^{4/3} 4\pi x^{2} dx,$$

$$E_{c} = -c_{c} \int \rho^{7/6} 4\pi x^{2} dx.$$

### III. NUMERICAL PROCEDURES

For a given atomic number Z, a family of solutions of the corrected TFD equation, corresponding to different degrees of compression of the element, is obtained by varying the slope of the potential function at the origin. Several parameters of the integration are determined directly by this initial slope, which we denote by  $\phi'_{0}$ . From the discussion of Eq. (11) we conclude that the electron density, and hence  $\phi''$ , vanishes for x less than

$$x_{\gamma} = (1/s)[1-(1-a^2s/Z)^{1/2}],$$

where

$$s = -2\phi_0' + 1/\pi^2 Z.$$

In starting the stepwise numerical integration, three values of  $\phi$  and  $\phi''$  are used, including those at the origin. If we therefore choose an initial interval  $h_{in}$  such that

$$2h_{in} < x_1, \tag{22}$$

then  $\emptyset$  is linear in this region, and it is trivial to generate the starting values. For practical reasons  $h_{in}$  is chosen as the largest interval which satisfies both Eq. (22) and the condition

$$h_{in} = 0.02/2^b,$$
 (23)

where b is an integer. This is done so that upon doubling the space interval a number of times (not necessarily b times)  $\phi$  is evaluated at convenient values of x.

In the earlier work, justification was presented for determining the modification factor, a, through an "equivalent Coulomb problem." In following this procedure we consider a number of electrons interacting with the charged nucleus but not at all with each other, even to the extent of providing a partial screening of the nuclear charge. Under these conditions we would define of through the equation

$$\mathbb{Z}\emptyset = (\mathbf{E}' - \mathbf{V})\mathbf{x}$$
.

Here, in contrast to the situation in the actual atomic problem, the potential energy distribution is known. We have

$$V = -Z/x. (24)$$

It can be established by direct substitution that for the Coulomb problem,  $\phi$  is linear throughout the distribution, or

$$\emptyset = 1 + x\emptyset'$$
.

In order for Eq. (24) to be satisfied with no additive constant, we must let the Fermi energy of the Coulomb problem be given by

$$\mathbf{E'} = \mathbf{Z} \phi'. \tag{25}$$

Correction of the region of integration in the quantum-number plane is based on the above value of the Fermi energy, where for  $\phi'$  we use the initial slope of the actual atomic problem, i.e.,  $\phi'_0$ . The outer boundary of the region defines a quantity  $\alpha$  through the relation

$$\alpha = (n_r + \ell + 1)$$
 outer boundary.

The Fermi energy given by Eq. (25) is the maximum eigenvalue in the Coulomb field. From the form of these eigenvalues we obtain

$$\alpha = (-\mathbb{Z}/2\phi_0').$$

In correcting the integration region for the Coulomb field we derive the expression for the modification factor,

$$a = \alpha - (\alpha \nu)^{1/2},$$

where v is the (generally non-integral) number of filled shells obtained as the solution of the equation

$$v^2 + 3v/2 + 2\alpha^{3/2}v^{1/2} - (3\alpha^2 - 1/2) = 0.$$

Thus the initial slope of  $\emptyset$  determines the inner density cutoff distance  $x_1$ , the initial interval of integration  $h_{in}$ , and the modification factor a.

The quantum- and correlation-corrected TFD equation is of the form

$$\phi'' = f(x, \phi), \tag{26}$$

a form which can be integrated simply and rapidly by a finite-difference method described by Hartree. In this method the approximation is made that

$$\Delta^2 \phi_0 = h^2 (\phi_0'' + \Delta^2 \phi_0'' / 12), \qquad (27)$$

the subscripted quantities here being associated with the point  $x=x_0$  to which the integration has progressed. In the usual notation,  $\Delta^2$  is the second difference operator, such that

$$\Delta^{2} \phi_{0} = (\phi_{1} - \phi_{0}) - (\phi_{0} - \phi_{-1})$$
$$= \phi_{1} - 2\phi_{0} + \phi_{-1},$$

and h is the existent integration interval.

To proceed in the integration an estimate is made of  $\Delta^2 \phi_0''$ , and from Eq. (27) we find  $\Delta^2 \phi_0$ . From the backward first difference and  $\Delta^2 \phi_0$  we can predict  $\phi$  at the next step. Eq. (26) then furnishes the predicted value of  $\phi''$ , from which the predicted  $\Delta^2 \phi_0''$  follows. This predicted value is compared with the original estimate to determine whether the integration is to be allowed to proceed to the next step, or whether it must be repeated with a revised estimate of  $\Delta^2 \phi_0''$ . The criterion for this decision is discussed below.

Some modification of the integration procedure seems advisable in the vicinity of  $x_1$ , where an abrupt change in  $\emptyset''$  occurs. A table of differences in  $\emptyset''$  of second-order and above reveals that the assumption that leads to Eq. (27), namely, that the terms involving differences higher than second-order can be neglected, is not too well justified

for x close to  $x_1$ . We can attempt to do a little better by adding one more term and writing

$$\Delta^{2} \phi_{0} = h^{2} (\phi_{0}'' + \Delta^{2} \phi_{0}'' / 12 - \Delta^{4} \phi_{0}'' / 240). \tag{28}$$

The extra term is retained for only a few integration steps for which the changes in  $\emptyset''$  are relatively large; in practice it is dropped upon reaching the point at which h has achieved its maximum value. We note that, for that portion of the integration in which Eq. (28) is utilized, it is necessary at each step to estimate  $\Delta \phi_0''$ ; but we can get an estimate of  $\phi''$  at the forward steps of sufficient accuracy to compute this difference by merely extending  $\phi$  linearly from the origin, thereby obtaining the arguments for Eq. (26).

The integration interval, starting at h<sub>in</sub>, is doubled on alternate steps until a certain maximum value is obtained, and then is kept constant out to the outer boundary of the charge distribution. This maximum value is selected by requiring that the precision in each integration be independent of Z, the precision being that of a chosen test run. A convenient check on the precision is furnished by the relative discrepancy between the total number of particles N which enters the boundary condition, Eq. (21), and the volume integral of the calculated electron density. It is thus apparent that there are two conditions on the integration. With h given, the criterion on proceeding to the next step in integrating the differential equation is that

$$\frac{\left|\frac{(\Delta^2 \phi_0'')}{\text{estimated}} - (\Delta^2 \phi_0'')\right|}{6\phi_0} \quad \text{predicted} \quad \text{h}^2 \leq 10^{-n},$$

where n is the number of significant digits carried in the calculation of  $\phi$ . This condition arises from requiring that an error in  $\Delta^2 \phi_0''$  ultimately cause an error in  $\phi_1$  of no more than 1/2 in the least significant digit. However, the precision of the integration, as measured by the

calculated number of particles, also depends upon h. We might expect this error to be dependent to a large degree upon the magnitudes of  $\Delta^2 \phi$  encountered in the integration, and results seem to bear this out. We require, then, as a rough measure of the error,

$$h_{\text{max}}^2 \phi_{\text{max}}'' \approx \text{constant},$$

and seek to estimate  $\phi''_{\max}$  as a function of Z and  $\phi'_{0}$ .

In the neglect of exchange and correlation effects we have

$$\phi'' = (4x/3\pi Z)(2Z\phi/x-a^2/x^2)^{3/2}$$
.

With Z and a given, the condition for a maximum of  $\phi''(x)$  is easily derived as

$$2a^2/x - Z(\phi-3\phi'x) = 0;$$

and since the maximum occurs at a small value of x, it is adequate for this discussion to put  $\phi = 1$  and to neglect  $3\phi'x$  in comparison with it. We then obtain the result that the maximum is at a position

$$x \approx 2a^2/Z$$

from which there follows

$$\phi_{\text{max}}'' \approx (3^{1/2}/\pi)Z/a$$
.

We then have the requirement

$$h_{\text{max}}^2 \approx \text{constant} \cdot (\pi/3^{1/2}) a/Z.$$

It is found, moreover, that a varies but slightly with Z and  $\phi_0'$ . We can treat it as a constant here. It is also found in our calculations that an interval  $h_{max} = 0.00125$  produces a respectably small error in number of particles of about four parts in  $10^6$  for  $Z = 5^4$ , with a modification factor of about 0.261. If  $h_{max}$  is chosen as the largest value obtained by doubling  $h_{in}$  subject to the condition

$$h_{mex} < 0.018/z^{1/2}$$
,

a fairly uniform error of a few parts in  $10^6$  results for all integrations, although for very small Z the error tends to be somewhat larger, say one part in  $10^5$ .

The outer boundary of the electron distribution,  $x_2$ , is determined by Eq. (21). We define a quantity

$$g = (x_0-h/2)(\phi_0-\phi_{-1})/h - (\phi_0+\phi_{-1})/2 + (Z-N)/Z,$$

which first becomes positive somewhere in the vicinity of  $x_2$ . At the integration step at which this occurs, a parabola is passed through the points  $\phi_{-2}$ ,  $\phi_{-1}$ , and  $\phi_0$ . We then have for this limited region the approximation

$$\phi = 0x^2 + \alpha x + C,$$

and the coefficients G, B, and C are evaluated under the condition that

$$\phi'' = \Delta^2 \phi/h^2.$$

Thus,

$$G = \phi''/2,$$

$$\beta = (\phi_0 - \phi_{-1})/h + hG - 2x_0$$

$$c = \emptyset_0 - 0x_0^2 - 0x_0.$$

The boundary condition becomes

$$0x_2^2 + 6x_2 + C = 20x_2^2 + 6x_2 + (Z-N)/Z$$

and hence

$$x_2 = [(ZC-Z+N)/ZG]^{1/2}.$$

Integrals yielding the total energies of the various forms and the total number of electrons are evaluated by the Simpson "1/3" rule, with boundary corrections at  $x_1$  and  $x_2$  computed by the trapezoidal rule.

### IV. RESULTS

It was pointed out in the Introduction that the quantum-corrected TFD equation yields atomic binding energies in good agreement with experimental values and with the results of Hartree-type calculations. It is of interest to know whether the agreement is retained when correlation energy is included. We also wish to ascertain the effect that inclusion of correlation has on the radii of the minimum-energy solutions. The pressure-compression curves on the TFD model suffer from this radius being too large for almost all elements, and correlation effects are known within the statistical theory to contract the electron cloud.

Table I presents summaries of minimum-energy solutions for a number of neutral atoms. We should mention here that the correlation energy for the low-Z elements is roughly twice that given by Clementi. 17

Table II compares the calculated total energies with Hartree-Fock-Slater non-relativistic values, and, for low-Z elements, with experimental

TABLE I MINIMUM-ENERGY SOLUTIONS

### A. THE INITIAL SLOPES, MODIFICATION FACTORS, AND INNER AND OUTER RADII

z -	<u>-ø'</u>	a —	<u>x</u> 1	<u>x</u> 5
2	1.662286185	.28930262	.021720161	3.1500
3	1.944982019	<b>.</b> 28290387	.013707739	3.3149
4	2.179156982	.27921366	.015083695	3.4274
5	2.372918675	.27665497	.0077987352	3.5175
6	2.548968232	.27480082	.0063976318	3.5925
7	2.705926337	.27334467	.0054165521	3.6525
8	2.852616885	.27218124	.0046931340	3.7075
10	3 <b>.</b> 108618420	.27036059	.0036973067	3.7925
20	4.063980433	.26564491	.0017783697	4.0387
<b>3</b> 0	4.745875374	.26359841	.0011645068	4.1787
40	5.289883861	.26226632	.00086374273	4.2637
50	5.752752881	.26133199	.00068564904	4.3387
60	6.158923349	.26062552	.00056803471	4.3893
70	6.520766586	.26006289	.00048462240	4,4306
80	6.852574996	.25960300	•00042243370	4.4729
90	7.157896787	•25921554	.00037429564	4.4972
100	7.440332217	.25888199	.00033593916	4.5293

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TABLE I -- Continued

B. THE VARIOUS FORMS OF ENERGY (a.u.)

Z -	Ef	Eq	Ep	E <sup>e</sup> _p	Eex	Ec	E —
2	2.2270	0.67061	-6.5965	1.5881	<b>-</b> 0.73537	-0.11198	-2.9582
3	5.9 <del>4</del> 21	1.5576	-17.415	<b>3.</b> 9161	-1.4135	-0.18777	-7.6005
4	11.962	2.8506	-34.693	7.4654	-2.2572	-0.27172	-14.944
5	20.611	4.5210	-59.182	12.345	<b>-3.251</b> 9	-0.36243	<b>-25.31</b> 9
6	32.149	6.5894	<b>-</b> 91 <b>.</b> 534	18.646	<b>-4.386</b> 7	-0.45896	-38.995
7	46.827	9.0707	-132.36	26.453	<b>-</b> 5.6544	-0.56071	<b>-</b> 56.225
8	64.846	12.030	-182.22	<b>35.83</b> 8	<b>-</b> 7 <b>.</b> 0479	-0.66715	-77.217
10	111.73	19.085	-310.64	59.605	-10.193	-0.89269	-131.31
20	602.82	80.184	-1625.0	292.06	-32.214	-2.2164	-684.37
<b>3</b> 0	1609.6	185.67	-4269.1	743.60	-63.280	<b>-3.</b> 78 <b>2</b> 9	-1797.3
40	3225.3	33 <sup>4</sup> .76	-8461.7	1445.6	-102.22	<b>-</b> 5.5 <b>32</b> 9	-3563.9
50	5523.5	529.12	-14378.	2422.7	-148.31	-7.4337	-6058.2
60	8566.6	771.39	-22167.	3696.1	-201.05	<b>-9.</b> 4648	-9343.2
70	12409.	1057.3	-31953.	5 <b>2</b> 83.9	-260.02	-11.611	-13474.
80	17100.	1386.9	-43849.	7202.5	-324.94	-13.861	-18498.
90	22683.	1764.7	-57964.	9466.8	-395.53	-16.206	-24461.
100	29200.	2186.9	-74392.	12090.	<del>-4</del> 71.58	-18.639	-31405.

TABLE II

COMPARISON OF CALCULATED AND EXPERIMENTAL TOTAL BINDING ENERGIES (a.u.)

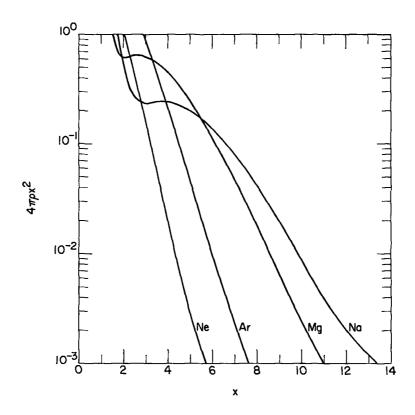
z <del>-</del>	-Ecalc	-E <sub>HFS</sub>	-E exp
2	2.9582	2.8779	2.9027
3	7.6005	7.2262	7.4761
4	14.944	14.255	14.665
5	25.319	24.079	24.652
6	<b>3</b> 8.995	37.079	37.846
7	56.225	53.587	5 <sup>4</sup> •598
8	77.217	73.938	75.092
10	131.31	127.48	
20	684.37	674.02	
30	1797.3	1773.6	
40	3563.9	3532.6	
50	6058.2	6014.6	
60	9343.2	9273.6	
70	13474.	13380.	
80	18498.	18395.	
90	24461.	24343.	
100	31405.	31264.	

values. <sup>18</sup> The Hartree-Fock-Slater results through Z = 30 were reported by Snow, et al, <sup>19</sup> as computed from the code published by Herman and Skillman, <sup>20</sup> while for larger Z the values were calculated by Cowan with a modification of the same code. In comparing with experiment, the binding energies on the corrected TFD model are seen to be not much worse than those calculated by the self-consistent field method, and in some cases are better. At high Z the two methods give energies differing by roughly one-half of one percent. It should be mentioned, however, that the agreement is slightly better on the model that includes the quantum correction but no correlation. This is especially true for lower Z. Correlation increases the discrepancies with experimental energies from about 2% to about 3% for atomic numbers 6, 7, and 8, but at high Z it causes a change in energy of only about 0.1%.

The radius of minimum-energy solutions as a function of atomic number is shown in Fig. 2; this radius is the "lattice constant" if the atoms are considered to be bound in a solid. Also shown are the corresponding TFD values computed by Thomas, 22 and spherical cell radii calculated from the observed normal crystal densities. The inclusion of correlation produces a cell radius which represents a somewhat better average to the experimental points in the variation with Z, but quite obviously the equation of state for many metals will not be greatly improved.

Although in equation of state calculations a zero boundary pressure solution is tacitly considered to represent an atom bound in a crystal of normal density, the calculated radii of such solutions actually support their interpretation as representing isolated atoms. One might object that, with reference again to Fig. 2, the calculated cell radius of some elements, notably most of the alkalis and alkaline earths, is <u>less</u> than the observed crystal radius. This result is not surprising for these elements, as can be seen from the sketch on page 35. Here are shown the radial distribution functions, as calculated by the self-consistent field code of Boyd, et al, <sup>23</sup> for the ground states of neon, sodium, magnesium, and argon. The long "tail" of the distribution, apparent especially for

Figure 2. Lattice constant of the elements.



sodium, and to a lesser extent for magnesium, is not obtained on the statistical model. The statistical density distributions have an abrupt cutoff, and thus much more closely resemble those of the rare gases.

The radii of the rare-gas atoms neon, argon, krypton, and xenon, computed on the present model agree closely with their crystal radii. This comparison is made in Table III, where the "experimental" values are computed from the experimental lattice constants given in the recent review article by Pollack. The rare gases are bound in crystals only by the very weak van der Waals forces; if we were to ascribe a finite radius to the isolated atom, it should be for the solid rare gases that such a radius would most nearly equal its crystal radius. Since the corrected statistical model predicts close to these values for the rare-gas atomic radii, it would appear that the correct interpretation of minimum-energy, or zero boundary pressure, solutions is as representing isolated atoms.

TABLE III

LATTICE CONSTANTS OF THE SOLID RARE GASES (a.u.)

Element	(x <sub>2</sub> ) <sub>calc</sub>	$(x_2)_{exp}$
Neon	<b>3.</b> 79	3. <i>3</i> 0
Argon	4.00	3.92
Krypton	4.23	4.17
Xenon	4.36	4.53

# APPENDIX

# A FORTRAN CODE TO INTEGRATE THE QUANTUM- AND CORRELATION-CORRECTED TFD EQUATION

The FORTRAN machine code listed here in "F4" language generates a single solution of the quantum- and correlation-corrected TFD equation for a given atomic number, initial slope of the potential function, and degree of ionization. The minimum-energy solutions were stressed in this report, and the code as actually used contains a feature that searches for the solutions possessing the lowest energy by adjusting  $\phi_0'$  and performing a series of integrations. However, there seems little virtue in complicating the present write-up by including a number of code statements that are unnecessary for the task to which a potential user may put the code.

The input data consists of any number of sets of Z,  $\phi'_0$ , degree of ionization, and a print flag that indicates whether the entire solution is to be printed, or whether summary information only is desired. Each set is entered by a data card, the layout of which is as follows:

Columns	<u>Data</u>
1 - 4	Atomic number Z
5 <b>-</b> 18	Initial slope $\phi_0'$
19 - 23	Degree of ionization
<b>5</b> /4	Print flag

All but the print flag are floating point numbers. For example, to obtain and print the complete solution corresponding to the free neutral lithium atom, one would prepare the following card:

Columns	<u>Data</u>
1 - 4	003.
5 <b>-</b> 18	-1.94498201900
19 - 23	0000.
24	1

If only summary information were desired, column 24 would contain a 0 punch.

As a further example, a solution corresponding to a compressed 0 ion of radius 2.5203 a.u. is obtained with the input

Columns	Data
1 - 4	008.
5 <b>-</b> 18	-2.84800000000
19 - 23	-002.
24	1

The output contains a listing of Z and  $\phi'_0$ , and the computed value of a. There follows, unless suppressed by the presence of a 0 in column  $2^{1}$  of the input card:  $x_1$ ,  $\rho(x_1)$ , and a tabulation of  $\phi$  and  $\rho$  for each x value. Immediately following the tabulation is the computed outer radius of the distribution,  $x_2$ , and the interpolated values of  $\phi(x_2)$  and  $\rho(x_2)$ . Also printed out are the total calculated number of electrons within the distribution, and the various energies. If the printing is suppressed, then  $x_1$ ,  $\rho(x_1)$ , and the table of  $\phi(x)$  and  $\rho(x)$  are not printed.

The code consists of a main program and a number of FUNCTION and SUBROUTINE subprograms. "Comment" cards make the purpose of the programs evident; no further explanation of their purpose is required here. However, an attempt will be made to clarify a few items that might

# prove puzzling:

In the SOURCE subprogram, note that after 10 attempts to find y by iteration, y being a solution of Eq. (15) of the text, the initial guess is changed; and 10 more attempts can be made. This can occur only once in each integration, where the density decreases below about 0.0123. At this point there becomes only one real root of the cubic equation, this root being approximately 0.13, whereas the iteration procedure utilizing the solution on the previous step of the integration as a first guess attempts to find a solution near y = -0.18.

Another item is an apparently extraneous integral that is calculated and never used. This is calculated through FLINT(8) in the integrand routine, and SUM(8) in the main routine. The integral is

$$\int_{\mathbf{x}_{1}}^{\mathbf{x}_{2}} \rho \mathbf{x}^{4} d\mathbf{x},$$

from which the diamagnetic susceptibility and other quantities of possible interest can be calculated if desired. It should be mentioned that, indeed, the susceptibility has been calculated on this model for the rare gases. The agreement with experiment is slightly better than on the uncorrected TFD model; but the latter values are already in quite good agreement, and the improvement is small.

In addition to the possible necessity of altering the Input-Output statements of the code, it may also be necessary to change the iteration criteria used in the AFUNCT, INTSEC, and SOURCE subprograms.

## SUBTYPE, FORTRAN

	С		MAIN PROGRAM TO INTEGRATE THOMAS-FERMI-DIRAC EQUATION	JN .	MAINMTFD	
	С		WITH QUANTUM AND CORRELATION CORRECTIONS		MAINMTFD	
00000			DIMENSION X(3), PHI(3), FNT(3,8), SUM(8), FNTTM(8)		MAINMTED	000002
00001			COMMON Z.PI.PISQRD.C1,C2,C3,TAU1,CSLOPE,HMAX,R,Y,PH	(X2,FLINT(8)	MAINMTFD	000002
	С		CONSTANTS		MAINMTFD	000002
00002			PI=3.14159265		MAINMTFD	000002
00003			PISQRD=PI*PI		MAINMTFD	000004
00004			CON1=3.0**(1.0/3.0)		MAINMTFO	000007
00005			CON2=PI •• (1.0/3.0)		MAINMTFD	000021
30000			CF=.3*PI*CON1*CON1*CON2		MAINMTFD	000033
00007			CEX=.75+CON1/CON2		MAINMTFD	000041
00008			CC=.0842		MAINMTED	000045
00009			C1=8.0/(3.0*PISQRD)**(2.0/3.0)		MAINMTFD	000047
00010			C2=.5/(P1SQRD)		MAINMTFD	000065
00010			C3=[49.0/36.0]+CC+CC+4.0/(3.0+PISQRD]++(4.0/3.0)		MAINMTFD	000070
					MAINMTED	000114
00012			C4=4.0=FI		MAINMTED	000117
00013	_		TAU1=2.0/(3.0+PI++5)++(1.0/3.0)		MAINMTFD	000117
	С	_	READ PROBLEM INPUT			000117
00014			READ 10,2, SLOPE, DEGION, IPRTFG		MAINMTFD	
00015		10	FORMAT(F4.0,F14.0,F5.0,I1)		MAINMTFD	000177
00016			CSLOPE=SLOPE		MAINHTFD	000177
00017			RHOCON=Z/C4		MAINMTFD	000201
	C		OBTAIN MEDIFICATION FACTOR		MAINMTFD	000201
00018		15	A=AFUNCT(SLOPE)		MAINMTFD	000204
	С		OBTAIN INNER RADIUS OF ELECTRON DISTRIBUTION		MAINMTFD	000204
00019			X1=X1FNCT(SLOPE,A)		MAINMTFD	000211
00020			PHIX1=1.0+SLOPE*X1		MAINMTFD	000220
	С		INITIAL GUESS FOR Y		MAINMTFD	000220
00021	•		Y=19		MAINMTFD	000224
CCOZZ			RHOX1=(RHOCON/X1)+SOURCE(X1,PHIX1,A)		MAINMTED	000226
30022	С		RESTORE PAPER AND PRINT PROBLEM DATA		MAINMTFD	000226
00023	•		PRINT 20		MAINMTFD	000243
00025		20	FURMAT(1H1)		MAINMTFD	000254
00025		20	PRINT 22.Z		MAINMIFD	000254
00025		2.2	FORMAT(7H Z= 15)		MAINMTFD	000272
00027		44	PRINT 24, SLOPE		MAINMTFD	000272
		٠,			MAINMTED	000310
00028		24	FORMAT(7H SLOPE= F16.10)		MAINMIFD	000310
00029			PRINT 26,A			000326
00030		26	FORMAT(7H A= F16.10)		MAINMIFD	000326
00031			IF(IPRTFG)261,36,261		MAINMTFD	000320
00032			PRINT 28,X1		MAINMTFD	
00033		28	FORMAT(7H X1= F16.10)		MAINMTFD	000346
00034			PRINT 30,RHOX1		MAINMTFD	000346
00035		30	FORMAT(7H RHUX)=F13.7///)		MAINMTED	000364
	С		PRINT COLUMN HEADINGS		MAINMTED	000364
00036			PRINT 35		MAINMTED	000364
00037		35	FURMAT(8H X,13H PHI,16H	RHO//)	MAINMTFD	000375
	С		COMPUTE INITIAL AND FINAL INTEGRATION INTERVALS		MAINHTFD	000375
00038		36	SPLIT=1.0		MAINMTFD	000375
00039			IF(SPLIT04/X1)45,45,50		MAINMTFO	000377
00040			SPLIT=SPLIT+2.0		MAINMTFD	000404
00041			GO TO 4C		MAINMTFD	000407
00042		50	H=.02/SPLII		MAINMIFD	000410
00042		-5	HMAX=.018/SQRT(Z)		MAINMTFD	000413
343	С		CALCULATE INITIAL VALUES FOR INTEGRATION		MAINMTFD	000413
	L		CHECOCHIE INTITAL AMENGS LOW THICOGRAPHON			

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C0044		DO 55 I=1,3	MAINMTFD	000417
00045		FLI=I	MAINMTFD	000420
00046		X(I)=(FLI-1.0)+H	MAINMTFD	000422
00047	55	PHI(I)=1.0+SLOPE+X(I)	MAINMTED	000426
00048		IF(IPRTFG)56,600,56	MAINMTFD	000420
	С	PRINT FIRST THREE POINTS		
00049			MAINMTED	000432
		PRINT 60,(X(1),PHI(1),1=1,3)	MAINMTFD	000434
OC050		FORMAT(2F12.8)	MAINMTFD	000460
00051	600	PHIM2=1.0	MAINMTED	000460
CG052		PH1M1=PH1(2)	MAINMTFD	000462
00053		PHIO=PHI(3)	MAINMTFD	000464
00054		PHDPM2=0.0	MAINMTFD	000466
00055		PHDPM1=0.0	MAINMTFD	000470
00056		PHIDPO=0.0		
CC057		X0=X(3)	MAINMTFD	000472
			MAINMTFD	000474
CC058		INDXD8=1	MAINMTFD	000476
00059		XM1=X1	MAINMTFD	000500
06000		DD 61 J=1,8	MAINMTFD	000502
CC061	61	SUM(J)=0.0	MAINMTED	000503
00062		CALL FLINTS(X1,PHIX1,RHOX1)	MAINMTFD	000505
00063		DO 62 J=1.8	MAINMTFD	000515
00064	6.2	FNT(2,J)=fL[NT(J)		
00004	C		MAINMTFD	000516
	-	ADVANCE ONE STEP IN X	MAINMTED	000516
00065	65	CALL INTSEC(PHIM2,PHIM1,PHIO,PHDPM2,PHDPM1,PHIDPO,H,XO,A)	MAINMTFD	000520
00066		IF(X0-X1)70,75,75	MAINMTED	000544
00067	70	IF(IPRTFG)71,72,71	MAINMTFD	000547
00068	71	PRINT 60.XO.PHIG	MAINMTED	000551
00069	72	INDXDE=-INDXDB	MAINMTFD	000574
0GU70		GU TO 65	MAINMIFE	000576
00071	75	RHO=RHOCON•PHIDPO/XO	MAINMTFD	000577
0.0072				
		IF(IPRTFG)76,813,76	MAINMIFD	000603
00073		PRINT 80, XO, PHIC, RHO	MAINMTFD	000605
00074		FORMAT(2F12.8,F17.7)	MAINMTED	000635
00075	810	CALL FLINTS(XO, PHIO, RHO)	MAINMTFD	000635
00076		DO 81 J=1,8	MAINMTFD	000645
00077		FNT(3,J)=FLINT(J)	MAINMTFD	000646
00078		SUM(J)=SUM(J)+.5*(FNT(2,J)+FNT(3,J))*(XC-XM1)	MAINMIFD	000650
00079	81	FNT(2,J)=FNT(3,J)	MAINMTFD	000661
00080	٠.	XM1=X0		
			MAINMTFD	000663
00081		INDXD8=-INDXD8	MAINMTFD	GOC665
00082		IF(INDXDB)65,65,85	MAINMTFD	000667
00083		CO 86 J=1,8	MAINMTFD	000671
00084	86	FNT(1,J)=FNT(3,J)	MAINMTFC	000672
00085		IF(H5*HMAX)90,94,94	MAINMTFD	000674
	C	DOUBLE INTEGRATION INTERVAL	MAINMTFD	000674
00086	90	H=Z.04H	MAINMTED	000701
0C087		PHIM1=PHIM2	MAINMTED	000704
00088		PHIM2=1.C		
			MAINMTFD	000766
00089		PHDPM1=PHDPM2	MAINMTFD	000710
00090		PHOPM2=0.0	MAINMTFD	000712
00091	91	CALL INTSEC(PHIM2,PHIM1,PHIO,PHDPM2,PHDPM1,PHIDPO,H,XO,A)	MAINMTFD	000714
00092		RHC=RHOCON*PHIDPC/XO	MAINMTFD	000740
00093		IF(IPRFFG)910.911.910	MAINMTED	000744
00094	910	PRINT 80, XO, PHIO, RHU	MAINMTFD	000746
00095		CALL FLINTS(XO, PHIO, RHU)	MAINMTED	000776
00096		DU 92 J=1,8	MAINMTFD	001006
00097	0.*			
	74	FNT(3, J) = FL IN((J)	MAINMTFD	001007
00098		INDXDB=-INDXDB	MAINHTFD	001011
00099		IF(INDXDB)93,93,935	MAINMTFD	001013
OC100	93	DO 931 J=1,8	MAINMTFD	G01015
66101	931	FNT(2,J)=FNT(3,J)	MAINMTED	CO1016
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		00.70.01	MAINMTFD	001020
00102		GD TO 91	MAINMTED	001021
00103		DO 936 J=1,8	MAINMTED	001022
00104	936	SUM(J)=SUM(J)+(H/3.0)+(FNT(1,J)+4.0+FNT(2,J)+FNT(3,J))	MAINMTED	001034
00105	_	GO TO 85	MAINMTED	001034
	C	INTEGRATION PROCEEDS WITH CONSTANT X INCREMENT	MAINMTED	001035
CC106	94	CO 116 I=2,3	MAINMIFD	001036
00107		CALL INTSEC(PHIM2, PHIM1, PHIO, PHDPM2, PHDPM1, PHIDPO, H, XO, A)		001036
	С	TEST IF AUXILIARY CONDITIONS ON INTEGRATION	MAINMTFD	001036
	С	ARE SATISFIEC. VIOLATION INDICATES INITIAL SLOPE	MAINMTFD	001036
	С	IS TOO LARGE NEGATIVELY	MAINMTED	001062
00108		YSQPLR=Y*Y+R	MAINMTFD	
00109		IF(YSCPLK)98,96,96	MAINMTFD	001066
00110	96	IF(Y+TAU1)98,97,97	MAINMTED	001070
00111	97	IF(TAU1-Y+2.0*SQRT(YSQPLR))98,10C,100	MAINMTFD	001073
00112	98	PRINT 99,XC	MAINMTFD	001103
00113		FORMAT(34H AUXILIARY CONDITION VIOLATED X= F12.8)	MAINMTFD	001121
C0114	• •	GO TO 5	MAINMTFD	001121
00115	100	RHQ=RHQCQN+PHIDPO/XO	MAINMTFD	001122
00116	-00	CALL FLINTS(XO, PHIO, RHO)	MAINMTFD	001126
00117		DO 101 J=1,8	MAINMTFD	001136
00118	101	FNT(1,J)=FLINT(J)	MAINMTFD	001137
00119	101	IF(IPRTFG)105,11C,105	MAINMTFD	001141
	105	PRINT 80.0X,PHIO,RHO	MAINMTFD	001143
OC120		TEST IF OUTER BOUNDARY IS REACHED	MAINMTFD	001143
	С	G=(XO5+H)*(PHIO-PHIM1)/H-(PHIO+PHIM1)/2.0+DEGION/Z	MAINMTED	001173
00121	110		MAINMTED	001215
00122		IF(G)116,120,120	MAINMTFD	001217
00123	116	CONTINUE	MAINMTFD	001220
00124		DO 117 J=1,8	MAINMTED	001221
OC125		SUM(J)=SUM(J)+(H/3.0)+(FNT(1,J)+4.0+FNT(2,J)+FNT(3,J))	MAINMTFD	001233
00126		FNT(1,J)=FNT(3,J)	MAINMTED	001235
OC127	117	FNTTM(J)=FNT(3,J)	MAINMTED	001237
00128		XTRM=XO	MAINMTFD	001241
00129		GO TO 94	PAINMTED	001241
	С	OBTAIN OUTER RADIUS OF DISTRIBUTION		001242
00130	120	X2=X2FNCT(PHIM2,PHIM1,PHIO,X0,H,DEGION)	MAINMTFD	001242
OC131		RHOX2=RHOCON+SOURCE(X2,PHIX2,A)/X2	MAINMTFD	001274
00132		IF(IPHTFG)124,122,124	MAINMTFD	001276
00133	122	PRINT 123,X2	MAINMTFD	001314
00134	123	FURMAT(7H X2= Fll.5)	MAINMTFD	
00135		GO TO 126	MAINMTFD	001314
00136	124	PRINT 125,x2,PHIX2,RHOX2	MAINMTFD	001315
00137	125	FORMAT(//F8.4.F16.8.F17.7)	MAINMTED	001345
	c	CALC END POINT CORRECTIONS TO INTEGRALS	MAINMTFD	001345
00138	126	CALL FLINTS(X2, PHIX2, RHOX2)	MAINMTFD	001345
00139		DO 13C J=1,8	MAINMTFD	001355
00140	130	SUM(J)=SUM(J)+.5*(X2-XTRM)*(FNTTM(J)+FLINT(J))	MAINMTED	001356
00140	c	CALC ENERGIES FROM INTEGRALS	MAINMTFD	001356
00141	•	EPRIME=(Z*PHIX2-DEGION)/X2-C2	MAINMTED	001367
00142		FLNUM=C4+SUM(1)	MAINMTFD	001375
00142 00143		EPN=-Z+C4+SUM(4)	MAINMTFD	001400
00144		EPEPRM=-Z+C4+SUM(5)	MAINMTFD	001404
		EF=CF*C4*SUM(3)	MAINMTFD	001410
00145			MAINMTFD	001414
00146		EEX=-CEX+C4+SUM(2)	MAINMTFD	001420
00147		EQ=A+A+C4+SUM(6)/2.0	MAINMTED	001426
00148		EC=-CC+C4+SUM(7)	MAINMTFD	001432
00149		EPE=.5*(EPEPRM-EPN+FLNUM*(EPRINE+C2))	MAINMTED	001443
00150		E=EF+EPN+EPE+EEX+EQ+EC	MAINMTED	001452
00151		PRINT 135, FLNUM	MAINMTED	001470
00152	135	FORMAT(//21H NUMBER OF ELECTRONS= F11.6)	MAINMTED	001470
QC153		PRINT 136, EF, EQ	MAINMTED	001513
00154	136	FORMAT(//5H EF= 1PE15.7.7H EQ= 1PE15.7)	PATIENTO	004,713

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CC155 00156	PRINI 137.EPN.EPE.EEX.EC 137 FORMAT(5H EPN= 1PE15.7,7H EPE= 1PE15.7,7h 1 7H EC= 1PE15.7)	H EEX= 1PE15.7,	MAINMTFD MAINMTFD	001513 001550
00157 00158 00159	PRINT 138.E 138 FORMAT(5H E= 1PE18.10//) GO TO 5		MAINMTED Mainmted Mainmted Mainmtec	001550 001550 001566 001566
00160	END		MAINMTFD	C01567

SUBPROGRAM .MAIN. - COMPILE FIME 00G005 SECS. - NO. BINARY CARDS 0C0000 - LENGTH (8)001267 WORDS (=(10)0C0695)

## SUBTYPE, FORTRAN

	С		COMPUTE MODIFICATION FACTOR FROM INITIAL SLOPE	AFNTMTFD	
	č		BY NEWTON-RAPHSON METHOD	AFNTMTFD	
00000	-		FUNCTION AFUNCT(SLCPE)	AFNTMTFD	000002
00001			COMMON Z,PI,PISQRD,C1,C2,C3,TAU1,CSLOPE,HMAX,R,Y,PHIX2,FLINT(8)	AFNTMTFD	000002
00002			CRIT=1.0E-14	AFNTMTFD	000002
			ALPHA=SQRT(2/(2.0+ABS(SLOPE)))	AFNTMTFD	000004
00003			RODT1=SQRT(ALPHA)	AFNTMTFD	000010
00004				AFNTMTFD	000012
00005			F1=3.0*ALPHA-ALPHA5	AFNTMTFD	000017
00006	_		F2=ALPHA+ROOT1	AFNTMTFD	000017
	С		INITIAL GUESS FOR NU	AFNIMIFD	000022
00007		_	FLNU=ALPHA5	AFNIMIFD	000025
00008		5	ROOT2=SQRT(FLNU)	AFNIMTED	000027
00009			F=FLNU*(FLNU*(FLNU+1.5)+2.0*F2*ROOT2-F1)	AFNTMTFD	000042
CCOIO			FPRIME=3.0*FLNU*(FLNU+1.0)+3.0*F2*R00T2-F1	AFNIMIFO	000057
00011			IF(FPRIME)15,10,15		000057
	С		ADJUST NU IF DERIVATIVE OF F IS ZERO	AFNIMIED	000057
00012		10	FLNU=.99999•FLNU	AFNTMTFD	
00013			GO TO 5	AFNIMIFD	000064
	c		CORRECT NU	AFNTMTFD	000064
00014	_	15	CURR=-F/FPRIME	AFNTMTFD	000065
00015			FLNU=FLNU+CORR	AFNTMTFD	000070
00015			ERROR=ABS(CORR/FLNU)	AFNTMTFD	000073
00010	С		TEST IF CRITERION IS MET	AFNTMTFD	000073
00017	·		IF(ERROR-CRIT)20,20,5	AFNIMIFD	000076
		2.0	AFUNCT=ALPHA-SQRT(FLNU)+ROOT1	AFNIMIFD	000101
00018		20		AFNTMTFD	000106
00019			RETURN	AFNTMTFD	000107
00020			END		

SUBPROGRAM AFUNCT - COMPILE TIME 000003 SECS. - NO. BINARY CARDS 000000 - LENGTH (8)000135 WORDS (=(10)000093)

## SUBTYPE.FORTRAN

00000 00001 00002 00003 00004	С	COMPUTE INNER RADIUS OF ELECTRON DISTRIBUTION FUNCTION X1FNCT(SLOPE,A) COMMON Z,PI,PISQRD,C1,C2,C3,TAU1,CSLOPE,HMAX,R,Y,PHIX2,FLINT(8) S=(1.0-PISQRD=Z=2.0+SLOPE)/(PISQRD=Z) X1FNCT=(1.0-SQRT(1.0-A+A+S/Z))/S RETURN END	X1FNMTFD X1FNMTFD X1FNMTFD X1FNMTFD X1FNMTFD X1FNMTFD X1FNMTFD	000002 000002 000002 000014 000026 000027
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SUBPROGRAM X1FNCT - COMPILE TIME 000002 SECS. - NO. BINARY CARDS 000000 - LENGTH (8)000067 WURDS (=(10)000055)

#### SUBTYPE, FORTRAN

00000 00001 00002 00003 00004 00005 00006 00007 00008	COMPUTE OUTER RADIUS OF ELECTRON DISTRIBUTION  AND INTERPOLATE FOR PHI AT OUTER BOUNDARY  FUNCTION X2FNCT(PHIM2,PHIM1,PHIO1,XO,++,DEGION)  COMMON Z,PI,PISQRD,C1,C2,C3,TAU1,CSLOPE,+MAX,R,Y,PHIX2,FLINT(8)  SCRIPA=-5=(PHIO-2.0=PHIM1+PHIM2)/(+++)  SCRIPB=(PHIO-PHIM1)/H-SCRIPA=(2.0=X0-H)  SCRIPB=(PHIO-X0=(SCRIPA=X0+SCRIPB))  X2=SQRT((SCRIPC-DEGION/Z)/SCRIPA)  PHIX2=X2=(SCRIPA=X2+SCRIPB)+SCRIPC  X2FNCT=X2  RETURN  END	X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD X2FNMTFD	000002 000002 000002 000014 000026 000034 000041 000047 000051
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SUBPROGRAM X2FNCT - COMPILE TIME 000002 SECS. - NO. BINARY CARDS 000000 - LENGTH (8)000115 HORDS (=(10)000077)

#### SUBTYPE, FORTRAN

	С		SUBROUTINE WHICH INTEGRATES SECOND ORDER DIFFERENTIAL EQUATION	INTGMTFD	
	С		BY HARTREE (STOERMER) METHOD, WITH A CORRECTION TERM	INTGMTFD	
	C		INVOLVING AN ESTINATE OF FOURTH DIFFERENCE IN PHI	INTGMIFD	
	С		DOUBLE PRIME WHEN THIS QUANTITY IS LARGE	INTGMTFD	
00000			SUBROUTINE INTSEC(PHIM2,PHIM1,PHIO,PHDPM2,PHDPM1,PHIDPO,H,XU,A)	INTGMTFD	000002
00001			COMMON Z,PI,PISQRD,C1,C2,C3,TAU1,CSLQPE,HMAX,R,Y,PH1X2,FLINT(8)	INTGMTFD	000002
00002			CRIT=1.0E-13	INTGMTFD	000002
00003			N=G	INTEMTED	000004
00004			XPI=XO+H	INTGMTFD	000006
	С		TEST IF TO APPLY CORRECTION	INTGMTFD	000006
00005			IF(H54HMAX)10,5,5	INTGMTFD	000011
00006		5	IF(PHOPM2)15,10,15	INTEMTED	000016
	С		CALCULATE FOURTH ORDER TERM	INTGMTFD	000016
00007		10	XP2=X0+2.0+H	INTGMTFC	000020
80000			ESPHP1=1.0+CSLOPE*XP1	INTGMTFD	000024
00009			ESPHP2=1.0+CSLOPE*XP2	INTGMTFD	000030
00010			EPDPP1=SOURCE(XP1,ESPHP1,A)	INTGMTFD	000034
00011			EPDPP2=SOURCE(XP2,ESPHP2,A)	INTGMTFD	000045
00012			FOURTH=(EPDPP2-4.0*EPDPP1+6.0*PHIOP0-4.0*PHDPM1+PHDPM2)/240.6	INTGMTFD	000056
00013			GO TO 20	INTGMTFD	000074
00014		15	FOURTH=0.0	INTGMTFD	000075
	С		ESTIMATE SECOND DIFFERENCE IN PHI DOUBLE PRIME	INTGMTFD	000075
00015		20	D2PDPE=PHIDPO-2.0*PHDPM1*PHDPM2	INTGMTFD	600077
	С		CALCULATE PREDICTED VALUE FOR COMPARISON	INTGMTFD	000077
00016		25	D2PHI=(H+H)+(PHIDPO+(D2PDPE/12.0)-FOUR(H)	INTGMTED	000104
00017			PHIPI=2.0 *PHIO-PHIM1+D2PHI	INTGMTFC	000115
00018			PHDPP1=SOURCE(XP1,PHIP1,A)	INTGMTFD	000122
00019			D2PDPP=PHDPP1-2.0*PHIOPO+PHDPM1	INTGMTFD	000133
	C		TEST FOR CONVERGENCE	INTGMTFD	000133
00020			ERROR = ABS((D2PDPP-D2PDPE)/PHIG) + H+H/6.0	INTGMTFC	000140
00021			IF(ERPOR-CRIT)45,45,30	INTGMTFD	000150
00022		30	D2PDPE=D2POPP	INIGMTFD	000153
00023			N=N+1	INTGMIFD	000155
06024			IF(N-10)25,35,35	INTGMTFC	000160
0C025			PRINT 40	INTGMTFD	000163
00026		40	FORMAT(30H INTEGRATION CRITERION NOT MET)	INTGMTFD	000174
00027			GO TO 45	INTGMTFD	000174
85000		45	PHIM2 = PHIM1	INTGMTFD	000175
00029			PHIM1=PHIO	INTGMTFD	000177
00030			PHIC=PHIP1	INTGMTFD	000201
00031			PHDPM2=PHDPM1	INTGMTFD	000203
00032			PHOPM1=PHIOPO	INTGMTFD	000205
00033			PHIDPO=PHDPP1	INTGMTFD	000207
00034			XO=XP1	INTGMTFD	000211
00035			RETURN	INTGMTFD	000213
00036			END	INTGMTFO	000214

SUBPROGRAM INTSEC - COMPILE TIME 000003 SECS. - NO. BINARY CARDS 000000 - LENGTH (8)000255 WORDS (=(10)000173)

## SUBTYPE, FORTRAN

	Ç	COMPUTE SOURCE FUNCTION, WHICH IS EQUATED TO	SRCEMTFO	
	С	SECOND DERIVATIVE OF PHI	SRCEMTFD	
00000		FUNCTION SOURCE(X,PHI,A)	SRCEMTED	000002
00001		COMMON Z,PI,PISQRD,C1,C2,C3,fAU1,CSLOPE,HMAX,R,Y,PHIX2,FLINT(8)	SRCEMTFD	000002
00002		CRIT=1.0E-14	SRCEMTFD	000002
00003		N=O	SRCEMTFD	000004
00004		NFLAG=0	SRCEMTFD	000006
00005		R=C1+(Z+PHI/X-A+A/(2.0+X+X)-C2)	SRCEMTFD	000010
	С	EVALUATE Y BY NEWTON-RAPHSON METHOD	SRCEMTFD	000010
00006		F2=TAUl+R-C3	SRCEMTFD	000027
CCCG7	5	F=Y*(Y*(Y+TAU1)+R)+F2	SRCEMTFD	000033
00008		FPRIME=Y=(3.G+Y+2.G+TAU1)+R	SRCEMTFD	000042
CC009		IF(FPRIME)15,10,15	SRCEMTFD	000053
00010	10	Y=.99999	SRCEMTFO	000055
CCJ11		GO TO 5	SRCEMTFD	000060
00012	15	CORR=-F/FPRIME	SRCEMTFD	000061
CC013		Y=Y+CORR	SRCEMTFD	000064
00014		EKROR=ABS(CORR/Y)	SRCEMTFD	000067
0C015		IF(ERROR-CR1T)50,50,20	SRCEMTED	000072
00016	20	N=N+1	SRCEMTFD	000075
00017		1F(NFLAG)35,25,35	SRCEMTFD	000100
00018	25	IF(N-10)5,30,30	SRCEMTED	000102
00019	30	NFLAG=1	SRCEMTED	000105
00020		Y=C.13	SRCEMTFD	000107
00021		N=O	. SRCEMTFD	000111
00022		GO TO 5	SRCEMTFD	000113
00023	35	IF(N-10)5.40.40	SRCEMTFD	000114
00024	4 Ç	PRINT 45	SRCEMTED	000117
00025	45	FORMAT(23H CRITERION UN Y NOT MET)	SRCEMTED	000130
00026	50	ROOT = SQRT(Y = Y + R)	SRCEMTFD	000130
00027		PSI=SQRT(Y+TAU1)	SRCEMTED	000134
00028		SOURCE=(.5*P1*X/Z)*(TAU1+PSI+ROOT)**3	SRCEMTFD	000150
C0029		RETURN	SRCEMTFD	000165
00030		END	SRCEMTFD	000166
				333233

SUBPROGRAM SOURCE - COMPILE TIME 000003 SECS. - NO. BINARY CARDS 00000C - LENGTH (8)000177 WORDS (=(10)000127)

# SUBTYPE, FORTRAN

	С	COMPUTE INTEGRANDS FOR ENERGY INTEGRALS AND TOTAL	FLNIMIFD	
	č	NUMBER OF PARTICLES CHECK	FLNTMTFD	
05050	•	SUBROUTINE FLINTS(X,PHI,RHO)	FLNTMTFD	000002
00001		COMMON Z,PI,PISQRD,C1,C2,C3,TAU1,CSLOPE,HMAX,R,Y,PHIX2,FLINT(B)	FLNTMTFD	000002
00002		RTRHO=RHO+*(1.0/3.0)	FLNTMTFD	000002
00003		FL (NT (6) = RHO	FLNTMTFD	000014
00004		FL INT (4) = RH0 = X	FLNTMTFD	000016
00005		FLINT(1)=FLINT(4)*X	FLNTMTFO	000021
00006		FLINT(8)=FLINT(1)*X*X	FLNTMTFD	000024
00000		FLINT(2)=FLINT(1)*RTRHU	FLNTMTFD	000030
00008		FLINT(3)=FLINT(2)*RTRHO	FLNTMTFD	000033
00000		FLINT(51=PHI#FLINT(4)	FLNTMTFO	000036
00010		FLINT(7)=FLINT(1)+SQRT(RTRHO)	FLN1MTFD	000041
00011		RETURN	FLNTMTFD	000045
00011		ENO	FLNTMTFD	000046
CGGIL		2.10		

SUBPROGRAM FLINTS - COMPILE TIME 000002 SECS. - NO. BINARY CARDS 000000 - LENGTH (8)000105 WORDS (=(10)000069)

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